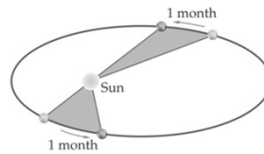


## Kepler's Three Laws of Planetary Motion

- 1<sup>st</sup> Law
  -
- 2<sup>nd</sup> Law
  -



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## Kepler's Three Laws of Planetary Motion

- 3<sup>rd</sup> Law
  - The square of the period of any planet is proportional to the cube of the planet's mean distance from the sun

Can be used for any  
object revolving  
around another.

T = period of the satellite **in seconds**

G = Gravitational Constant ( $6.67 \times 10^{-11}$ )

M = Mass of the object that is being orbited

r = distance between the center of the planet and the center  
of the sun

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## Kepler's Three Laws of Planetary Motion

- 3<sup>rd</sup> Law
  - For any objects orbiting the same planet or star:

Earth's Period around the sun = 365.25 days

Average distance from the sun to the Earth =  $1.5 \times 10^{11}$  m  
or 1 AU

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## Kepler's Three Laws of Planetary Motion

- 3<sup>rd</sup> Law (Example)
  - If it takes 686.95 days for Mars to revolve around the sun, what is its mean distance from the sun?

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## Newton's Universal Law of Gravity

- Any two objects of mass,  $m_1$  and  $m_2$  are accelerated towards each other by a force due to gravity.

$$G = 6.67 \times 10^{-11} \text{ N}\cdot\text{m}^2 / \text{kg}^2$$

$$M_E = 5.98 \times 10^{24} \text{ kg}$$

$$R_E = 6.38 \times 10^6 \text{ m}$$

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## Newton's Universal Law of Gravity

- For any object of mass,  $m$ , that is a certain distance from the surface of the Earth.

$$G = 6.67 \times 10^{-11} \text{ N}\cdot\text{m}^2 / \text{kg}^2$$

$$M_E = 5.98 \times 10^{24} \text{ kg}$$

$$R_E = 6.38 \times 10^6 \text{ m}$$

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## Gravitational Field

- The acceleration felt on a mass due to a gravitational force
- In general, the acceleration due to gravity is:
- 

where  $r$  is the distance from the center of the earth

- For any distance above the earth,  
 $r = R_E + h$

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## Gravitational Field

- How far above the surface of the earth must you be to have an acceleration due to gravity that is 85% of the gravity at the surface?

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